Introduction to the Standard Model William and Mary PHYS 771 Spring 2014

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Class information, including syllabus and homework assignments can be found at http://ntc0.lbl.gov/~walkloud/wm/courses/PHYS_771/

Homework Assignment 1

1. [5 pts.] We are primarily using the "mostly minus" metric, $g_{\mu\nu} = \text{diag}(1, -1, -1, -1)$. With this metric, the field strength tensor for a classical electromagnetic field is

$$F_{\mu\nu} = \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & -B_3 & B_2 \\ -E_2 & B_3 & 0 & -B_1 \\ -E_3 & -B_2 & B_1 & 0 \end{pmatrix}$$
 (1)

which can be compactly expressed as $F_{0i} = E_i$ and $F_{ij} = -\epsilon_{ijk}B_k$.

Solutions: It makes sense to solve (b) first, then (a). To begin, we note, in the mostly plus metric,

$$A^{\mu} = (\phi, \vec{A})^{T}, \qquad \partial_{\mu} = \left(\frac{\partial}{\partial t}, \vec{\nabla}\right)^{T}, A_{\mu} = (-\phi, \vec{A})^{T}, \qquad \partial^{\mu} = \left(-\frac{\partial}{\partial t}, \vec{\nabla}\right)^{T}.$$

From here, we can determine:

(b) Express the space-time F_{0i} and space-space F_{ij} components in terms of E_i and B_i .

[3 pts.] Solution:

$$F_{0i} = \partial_0 A_i - \partial_i A_0$$

$$= \frac{\partial}{\partial t} A_i - \nabla_i (-\phi)$$

$$= -E_i$$

And the space-space components are

$$F_{ij} = \nabla_i A_j - \nabla_j A_i$$
$$= \epsilon_{ijk} B_k$$

(a) Beginning with $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$, derive the form of $F_{\mu\nu}$ if we work with the "mostly plus" metric, $g_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$.

[2 pts.] Solution: From (b), we can immediately read off

$$F_{\mu\nu} = \begin{pmatrix} 0 & -E_1 & -E_2 & -E_3 \\ E_1 & 0 & B_3 & -B_2 \\ E_2 & -B_3 & 0 & B_1 \\ E_3 & B_2 & -B_1 & 0 \end{pmatrix}$$
 (2)

2. [10 pts.] We discussed using the covariant derivative to construct the field strength tensors for gauge theories, $igF_{\mu\nu} = [D_{\mu}, D_{\nu}]$. Suppose we have a fermion that is a doublet that transforms as

$$\psi(x) = \begin{pmatrix} \psi_1(x) \\ \psi_2(x) \end{pmatrix}, \quad \psi(x) \to e^{i\alpha^a(x)t^a} \psi(x), \quad \text{with } t^a = \frac{\sigma^a}{2}, \quad \sigma_a = \text{Pauli matrices}$$
 (3)

such that the covariant derivative is

$$D_{\mu} = \partial_{\mu} + igA_{\mu}(x) \qquad A_{\mu}(x) = t^{a}A_{\mu}^{a}(x) \tag{4}$$

(a) Derive the field strength tensor. You may find it useful to determine the components $igF^a_{\mu\nu}=[D_\mu,D_\nu]^a$ instead of $F_{\mu\nu}=F^a_{\mu\nu}t^a$. [5 pts.] Solution:

$$[D_{\mu}, D_{\mu}] = [\partial_{\mu} + igA_{\mu}, \partial_{\nu} + igA_{\nu}]$$

= $ig(\partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} + ig[A_{\mu}, A_{\nu}])$

The commutator of A-fields is

$$[A_{\mu}, A_{\nu}] = A^{a}_{\mu} A^{b}_{\nu} [t^{a}, t^{b}]$$
$$= A^{a}_{\mu} A^{b}_{\nu} i \epsilon^{abc} t^{c}$$

and so the field strength tensor is

$$F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} - g\epsilon^{abc}t^{c}A_{\mu}^{a}A_{\nu}^{b} \,.$$

We can project onto the a-th component using $tr[t^at^b] = \delta^{ab}/2$, and after relabeling dummy indices, we can write

$$F^{a}_{\mu\nu} \equiv 2\text{tr}[t^{a}F_{\mu\nu}] = \partial_{\mu}A^{a}_{\nu} - \partial_{\nu}A^{a}_{\mu} - g\epsilon^{abc}A^{b}_{\mu}A^{c}_{\nu}.$$

(b) In terms of the A^a fields, what is the form of the Lagrangian

$$\mathcal{L} = -\frac{1}{2} \text{tr}[F_{\mu\nu}^2] = -\frac{1}{4} (F_{\mu\nu}^a)^2 = ?$$
 (5)

[5 pts.] Solution:

$$(F_{\mu\nu}^a)^2 \equiv F_{\mu\nu}^a F^{a,\mu\nu}$$

$$= (\partial_\mu A_\nu^a - \partial_\nu A_\mu^a - g\epsilon^{abc} A_\mu^b A_\nu^c)(\partial^\mu A^{a,\nu} - \partial^\nu A^{a,\mu} - g\epsilon^{ade} A^{d,\mu} A^{e,\nu})$$

Relabeling the dummy indices, using the symmetry properties of ϵ^{abc} and the equality $\epsilon^{abc}\epsilon^{ade} = \delta^{bd}\delta^{ce} - \delta^{be}\delta^{cd}$, we find

$$\mathcal{L} = -\frac{1}{4} \left[(\partial_{\mu} A^{a}_{\nu} - \partial_{\nu} A^{a}_{\mu}) (\partial^{\mu} A^{a,\nu} - \partial^{\nu} A^{a,\mu}) - 4g \epsilon^{abc} A^{b}_{\mu} A^{c}_{\nu} \partial^{\mu} A^{a,\nu} + g^{2} (A^{a}_{\mu} A^{a,\mu} A^{b}_{\nu} A^{b,\nu} - A^{a}_{\mu} A^{a,\nu} A^{b,\mu} A^{b}_{\nu}) \right]$$

- 3. [5 pts.] For a classic electromagnetic field, Eq. (1),
 - (a) What is $F_{\mu\nu}F^{\mu\nu} = ?$ [2 pts.] Solution:

$$F_{\mu\nu}F^{\mu\nu} = F_{0i}F^{0i} + F_{i0}F^{i0} + F_{ij}F^{ij}$$

= $-2E_i^2 - \epsilon_{ijk}B_k(-\epsilon^{ijl}B_l)$
= $2(\mathbf{B}^2 - \mathbf{E}^2)$

(b) What is $\epsilon^{\mu\nu\rho\sigma}F_{\mu\nu}F_{\rho\sigma} = ?$ (with the convention $\epsilon^{0123} = +1$)
[3 pts.] Solution: With this convention, we can replace $\epsilon^{0ijk} = \epsilon^{ijk}$, $\epsilon^{i0jk} = -\epsilon^{ijk}$ etc. We then have

$$\epsilon^{\mu\nu\rho\sigma}F_{\mu\nu}F_{\rho\sigma} = \epsilon^{ijk}(F_{0i}F_{jk} - F_{i0}F_{jk} + F_{ij}F_{0k} - F_{ij}F_{k0})$$

$$= 2\epsilon^{ijk}(F_{0i}F_{jk} + F_{ij}F_{0k})$$

$$= 4\epsilon^{ijk}F_{0i}F_{jk}$$

$$= 4E_{i}\epsilon^{ijk}(-\epsilon^{jkh}B_{h})$$

$$= -8\mathbf{E} \cdot \mathbf{B}$$

- 4. [5 pts.] For $U(\lambda) = e^{i\lambda\alpha_a X_a}$ where X_a are the generators of a Lie Algebra,
 - (a) show $U(\lambda_1)U(\lambda_2) = U(\lambda_1 + \lambda_2)$ Solution: The crucial step is to realize $[\alpha_a X_a, \alpha_b X_b] = 0$ which is easy to show

$$\begin{split} [\alpha_a X_a, \alpha_b X_b] &= \alpha_a \alpha_b [X_a, X_b] \\ &= \left(\frac{1}{2} \{\alpha_a, \alpha_b\} + \frac{1}{2} [\alpha_a, \alpha_b] \right) [X_a, X_b] \\ &= \frac{1}{2} \{\alpha_a, \alpha_b\} [X_a, X_b] \\ &= 0 \end{split}$$

which all follows from symmetry/anti-symmetry. The vanishing commutation relation implies

$$e^{i\lambda_1\alpha_a X_a}e^{i\lambda_2\alpha_a X_a} = e^{i(\lambda_1 + \lambda_2)\alpha_a X_a}$$

and hence $U(\lambda_1)U(\lambda_2) = U(\lambda_1 + \lambda_2)$.

5. [5 pts.] For SU(2), what is the matrix form of the generators Solution: For any representation, we know the "ladder" operators span the space

$$J_{\pm}|jm\rangle = c_{im}^{\pm}|jm\pm 1\rangle\,, \qquad c_{im}^{\pm} = \sqrt{j(j+1) - m(m\pm 1)}$$

and that

$$[J_+, J_-] = 2J_3 , J_{\pm} = J_1 \pm iJ_2 .$$

These ladder operators can be easily constructed in a given representation by beginning with the lowest/maximum state, and raising/lowering it. The above relations can then be used to construct the generators.

(a) for the j = 1 representation? [3 pts.] Solution:

$$J_{+} = \begin{pmatrix} 0 & \sqrt{2} & 0 \\ 0 & 0 & \sqrt{2} \\ 0 & 0 & 0 \end{pmatrix} , \qquad \qquad J_{-} = \begin{pmatrix} 0 & 0 & 0 \\ \sqrt{2} & 0 & 0 \\ 0 & \sqrt{2} & 0 \end{pmatrix} ,$$

from which we get

$$J_3 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & -1 \end{pmatrix}, \quad J_1 = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \quad J_2 = \frac{i}{\sqrt{2}} \begin{pmatrix} 0 & -1 & 0 \\ 1 & 0 & -1 \\ 0 & 1 & 0 \end{pmatrix}.$$

(b) for the j = 3/2 representation? [2 pts.] Solution:

$$J_{+} = \begin{pmatrix} 0 & \sqrt{3} & 0 & 0 \\ 0 & 0 & 2 & 0 \\ 0 & 0 & 0 & \sqrt{3} \\ 0 & 0 & 0 & 0 \end{pmatrix}, \qquad J_{-} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ \sqrt{3} & 0 & 0 & 0 \\ 0 & 2 & 0 & 0 \\ 0 & 0 & \sqrt{3} & 0 \end{pmatrix},$$

from which we get

$$J_3 = \frac{1}{2} \begin{pmatrix} 3 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -3 \end{pmatrix}, \quad J_1 = \frac{1}{2} \begin{pmatrix} 0 & \sqrt{3} & 0 & 0 \\ \sqrt{3} & 0 & 2 & 0 \\ 0 & 2 & 0 & \sqrt{3} \\ 0 & 0 & \sqrt{3} & 0 \end{pmatrix}, \quad J_2 = \frac{i}{2} \begin{pmatrix} 0 & -\sqrt{3} & 0 & 0 \\ \sqrt{3} & 0 & -2 & 0 \\ 0 & 2 & 0 & -\sqrt{3} \\ 0 & 0 & \sqrt{3} & 0 \end{pmatrix}.$$

6. [10 pts.] Dirac algebra. In any representation, the Dirac matrices satisfy the algebra (in 4 dimensions)

$$\{\gamma^{\mu}, \gamma^{\nu}\} = 2g^{\mu\nu} \times \mathbb{1}_{4\times 4}. \tag{6}$$

In class, we defined the Dirac matrices in the "Dirac Basis", for which

$$\gamma_D^0 = \begin{pmatrix} \mathbb{1}_{2\times 2} & 0\\ 0 & -\mathbb{1}_{2\times 2} \end{pmatrix}, \quad \gamma_D^i = \begin{pmatrix} 0 & \sigma^i\\ -\sigma^i & 0 \end{pmatrix}, \quad \gamma_D^5 = \begin{pmatrix} 0 & \mathbb{1}_{2\times 2}\\ \mathbb{1}_{2\times 2} & 0 \end{pmatrix}, \quad \gamma^5 \equiv i\gamma^0\gamma^1\gamma^2\gamma^3.$$

$$(7)$$

Another useful and very common basis is the "chiral basis" (or Weyl basis) in which

$$\gamma_{\chi}^{0} = \begin{pmatrix} 0 & \mathbb{1}_{2 \times 2} \\ \mathbb{1}_{2 \times 2} & 0 \end{pmatrix}, \qquad \gamma_{\chi}^{i} = \begin{pmatrix} 0 & \sigma^{i} \\ -\sigma^{i} & 0 \end{pmatrix}, \qquad \gamma_{\chi}^{5} = \begin{pmatrix} -\mathbb{1}_{2 \times 2} & 0 \\ 0 & \mathbb{1}_{2 \times 2} \end{pmatrix}, \tag{8}$$

(a) Determine the similarity transformation which converts from the Dirac to chiral basis

$$\gamma_{\chi} = S\gamma_D S^{-1} \qquad S = ? \tag{9}$$

[4 pts.] Solution: This is simply a matter of diagonalizing γ_D^5 and using the eigenvectors to construct the rotation matrix. The only trick is to make sure we preserve the sign convention of the $\gamma_D^i = \gamma_\chi^i$ matrices. One finds

$$S = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & -1 \\ 1 & 1 \end{pmatrix}$$

(b) What is the similarity transformation that transforms from the chiral to Dirac basis?

[2 pts.] Solution: This is simply given by S^{-1} , as $\gamma_D^5 = S^{-1}S\gamma_D^5S^{-1}S$:

$$S^{-1} = S^T = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & 1 \\ -1 & 1 \end{pmatrix}$$

(c) In both the Dirac and chiral basis, in terms of the spinor components, what are

$$\psi_{\pm} = \frac{1 \pm \gamma^0}{2} \psi = ? \tag{10}$$

[2 pts.] Solution: If we write $\psi^T = (\psi_1, \psi_2, \psi_3, \psi_4)$ then we have

Dirac:
$$\psi_{+} = \begin{pmatrix} \psi_{1} \\ \psi_{2} \\ 0 \\ 0 \end{pmatrix}$$
 $\psi_{-} = \begin{pmatrix} 0 \\ 0 \\ \psi_{3} \\ \psi_{4} \end{pmatrix}$

Chiral: $\psi_{+} = \frac{1}{2} \begin{pmatrix} \psi_{1} + \psi_{3} \\ \psi_{2} + \psi_{4} \\ \psi_{1} + \psi_{3} \\ \psi_{2} + \psi_{4} \end{pmatrix}$ $\psi_{-} = \frac{1}{2} \begin{pmatrix} \psi_{1} - \psi_{3} \\ \psi_{2} - \psi_{4} \\ -\psi_{1} + \psi_{3} \\ -\psi_{2} + \psi_{4} \end{pmatrix}$

(d) In both the Dirac and chiral basis, in terms of the spinor components, what are

$$\psi_R = \frac{1+\gamma^5}{2}\psi = ?$$

$$\psi_L = \frac{1-\gamma^5}{2}\psi = ?$$
(11)

[2 pts.] Solution:

Dirac:
$$\psi_{R} = \frac{1}{2} \begin{pmatrix} \psi_{1} + \psi_{3} \\ \psi_{2} + \psi_{4} \\ \psi_{1} + \psi_{3} \\ \psi_{2} + \psi_{4} \end{pmatrix}$$
 $\psi_{L} = \frac{1}{2} \begin{pmatrix} \psi_{1} - \psi_{3} \\ \psi_{2} - \psi_{4} \\ -\psi_{1} + \psi_{3} \\ -\psi_{2} + \psi_{4} \end{pmatrix}$
Chiral: $\psi_{R} = \begin{pmatrix} 0 \\ 0 \\ \psi_{3} \\ \psi_{4} \end{pmatrix}$ $\psi_{L} = \begin{pmatrix} \psi_{1} \\ \psi_{2} \\ 0 \\ 0 \end{pmatrix}$

7. [35 pts.] In class, we discussed the g-factor for the electron and the nucleons. We saw in general, the elastic electromagnetic structure of a fermion, with parity conserving interactions, can be expressed as

$$\bar{u}(p')\Gamma^{\mu}(p',p)u(p) = \bar{u}(p')\left[\gamma^{\mu}F_1(q^2) + \frac{i\sigma^{\mu\nu}q_{\nu}}{2m}F_2(q^2)\right]u(p), \quad q = p' - p, \quad (12)$$

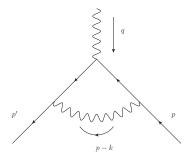


FIG. 1: The Feynman diagram used to compute g-2 of a point fermion.

where u(p) is an on-shell fermion spinor which satisfies pu(p) = mu(p) and $\sigma^{\mu\nu} \equiv \frac{i}{2}[\gamma^{\mu}, \gamma^{\nu}]$. This is the "elastic" structure, because $\bar{u}(p')$ also represents an on-shell fermion satisfying $\bar{u}(p)p = \bar{u}(p)m$.

In the case of the electron (point-like fermion) we saw the Dirac equation gives $g = 2 + \frac{\alpha_{f.s.}}{\pi}$. We noted that for the nucleons, $g_p \simeq 5.58$ and $g_n = -3.83$ so that the nucleons are not perturbatively close to point-like fermions, one indication they have interesting internal structure. We commented in class that $g = 2[F_1(0) + F_2(0)]$ and so for the electron, $F_2(0) = \frac{\alpha_{f.s.}}{2\pi}$.

Perform this classic QED calculation, using tools you have been learning in QFT, see Fig. 1. This calculation is so classic, you can easily find the solution in the literature. I strongly encourage you to attempt it on your own, before resorting external sources, peers, books, etc.. The key to successfully performing this calculation is to realize you isolate the contribution which is proportional to $\bar{u}(p')\sigma^{\mu\nu}q_{\nu}u(p)$. It turns out, this contribution to the diagram in Fig. 1 is free of both Ultraviolet (UV) $(q_E^2 \to \infty)$ and Infrared (IR) $(q_E^2 \to 0)$ singularities (where q_E is the Euclidean four-momentum obtained after Wick rotation of the momentum integral). To this end, recall the Gordon Identity which can be used to relate $\bar{u}(p')(p'+p)^{\mu}u(p)$ to $\bar{u}(p')\sigma^{\mu\nu}q_{\nu}u(p)$.

- (a) Compute g-2 for the electron [20 pts.] Solution: See attached hand notes.
- (b) Using just the requirements we have of our QFT, QED (renormalizable, gauge-invariant, Lorentz invariant QFT in 4 space-time dimensions) why should you know ahead of time that the contribution to g-2 is free of both UV and IR singularities?

[5 pts.] Solution: UV See attached hand notes.

[10 pts.] Solution: IR See attached hand notes.

The most general current between on-shell spinors is $\bar{\mathcal{U}}(p')$ $\bar{\mathcal{V}}_{\mu}(p',p)\mathcal{V}_{\mu}(p)$

The gyromagnetiz coefficient is defined as

Cument conservation gives F₁(o) = 1

We are interested in computing the correction, which is the coefficient of iturque 2m

To determine this quantity, we can couple the electron to a classical electromagnetic field with

The matrix element for an electron interacting with This external field is then given at leading order by

Re, p' | i A | ep> = U(p') (-i (-ê) Yn U(p) Aci, ê U(p) = -e U(p))

= -i e U(p') Yn U(p) Aci.

We want to compute the "radiative correction"

The factor of -ie and A'd will be the same for the one loop correction. So we can consider the correction to $\Gamma_{\mu}(p',p) = V_{\mu} + O(e^2)$

P' KT P-K

k' = k+9

(ep/ | i A lep) = -ieAd. \(\overline{\mu(p')} \int_{(2\overline{\gamma})}^{4k} \((-ie) \forall p \frac{-igp\sigma}{(p-k)^2 + i\tilde{\gamma}} \frac{i}{k'-m+i\tilde{\gamma}} \) \(\frac{i}{k-m+i\tilde{\gamma}} \) \(\frac{i}{k-m+i\tilde{\gamma}} \)

We can simplify the fermion propagators, \frac{1}{k-mtie} = \frac{k+m}{k^2-m^2 + i \varepsilon} and just look at the correction to In

=> $\overline{u}(p') \delta \Gamma_{n} u(p) = (-ie)^{2}(i)^{3}(-) \int \frac{d^{4}k}{(2\pi)^{4}} \frac{\overline{u}(p') \left[8^{p}(k'+m) \chi_{n}(k+m) \delta_{p}\right] u(p)}{\left[(k-p)^{2}+ie\right] \left[k'^{2}-m'^{2}+ie\right] \left[k'^{2}-m'^{2}+ie\right]}$ $= (-)^{3}(i)^{5} e^{2} \int \frac{d^{4}k}{(2\pi)^{4}} \frac{\overline{u}(p') \left[8^{p}k' \chi_{n}k' \chi_{p} + m' (8^{p}k' \chi_{n} \chi_{p} + 8^{p} \chi_{n} \chi_{p}) + m^{2} \delta^{p} \chi_{n} \chi_{p}}{\left[(k-p)^{2}+ie\right] \left[k'^{2}-m'^{2}+ie\right] \left[k'^{2}-m'^{2}+ie\right]} \times u(p)$

We can use the 4-d identities $Y^p Y_{\sigma} Y_{\nu} Y_{\nu} Y_{\rho} = -2 Y_{\nu} Y_{\nu} Y_{\sigma}$ $Y^p Y_{\nu} Y_{\nu} Y_{\rho} = 49 \mu \nu$ $Y^p Y_{\nu} Y_{\rho} = -2 Y_{\nu}$

 $\overline{u}(p') \, \delta \, \overline{u}(p) = -i \, e^2 \, \int d^4 t \, \overline{u}(p') \, \left[-2 \, k \, V_{\mu} t \, k' + 4 \, m \left(k_{\mu} t \, k'_{\mu} \right) - 2 \, m^2 \, V_{\mu} \right] \, u(p)} \, t_{\mu} = k \, \int \overline{u}(p') \, \left[-2 \, k \, V_{\mu} t \, k' + 4 \, m \left(k_{\mu} t \, k'_{\mu} \right) - 2 \, m^2 \, V_{\mu} \right] \, u(p)} \, d^2 t \, d^2$

Note: we have not included a photon wass (for IR regularization) nor have we worried about didimensional properties of Directlyebra. This is from the benefit of hindsight that the piece of this vertex we have are interested in is free of both IR and UV divergences.

To proceed with the integral, we can either combine the denominators with the Feynman Parameter trick

$$\frac{1}{a_1 a_2 - a_n} = \int_0^1 dx_1 dx_2 \cdot dx_n \cdot S(1 - \frac{g}{i = 1} x_i) \frac{(n-1)!}{\left[x_1 A_1 + x_2 A_2 + \cdots + x_n A_n\right]^n}$$

Dr we can use the Schwinger Integral trick $\frac{1}{k^2 - m^2 + i\epsilon} = \frac{1}{i} \int_0^\infty dz \, e^{iz(k^2 - m^2 + i\epsilon)}$

For fun, lets first solve with the Schwinger Integral trick.

We can shift the order of integration and look for a simple momentum integration variable. Completing the square, we find the argument of the exponential can be written

$$\tilde{L} \left\{ Z_{123} L^2 - m^2 \left(Z_2 + Z_3 - Z_1 + \frac{Z_1^2}{Z_{123}} \right) - q^2 \left(\frac{Z_1^2}{Z_{123}} - Z_2 \right) + \frac{2Z_1Z_2}{Z_{123}} P - q^2 \right\}$$
with $L = k - \frac{Z_1}{Z_{123}} P + \frac{Z_2}{Z_{123}} q$, $Z_{123} = Z_1 + Z_2 + Z_3$

For on-shell electrons, we also have

$$2 p \cdot q = (p+q)^{2} - p^{2} - q^{2}$$

$$= p^{2} - p^{2} - q^{2}$$

$$= m^{2} - m^{2} - q^{2}$$

Which allows us to reduce the argument to

$${\stackrel{\circ}{L}} \left\{ \begin{array}{c} {\frac{1}{2}} & {\frac{1}{2}}$$

Assuming we can shift integration order, we have

$$-2e^{2}\int_{0}^{\infty}dz_{1}dz_{2}dz_{3}\int_{0}^{4}t_{1} \overline{u}(p')\left[k_{1}y_{1}k'_{1}+m^{2}y_{1}-2m(k_{1}+k'_{1})\right]u(p)$$

$$\times \exp\left\{i\left[\overline{Z}_{123}l^{2}-\frac{\overline{Z}_{2}\overline{Z}_{3}}{\overline{Z}_{123}}q^{2}-\frac{(\overline{Z}_{2}+\overline{Z}_{3})^{2}}{\overline{Z}_{123}}m^{2}\right]\right\}$$

So, we now shift the momentum integration variable

$$= -2e^{2}\int_{0}^{\infty} dz_{1}dz_{2}dz_{3} \exp \left\{ i \left[\frac{Z_{123}l^{2} - \frac{Z_{2}Z_{3}}{Z_{123}}q^{2} - \frac{(Z_{2}rZ_{3})^{2}}{Z_{113}}m^{2} \right] \right\}$$

$$\int_{0}^{dAl} \times \tilde{\mathcal{U}}(P^{1}) \left[\left(\frac{1}{2} + \frac{Z_{1}}{Z_{123}} \cancel{X} - \frac{Z_{1}}{Z_{123}} \cancel{X} \right) \cancel{\mathcal{V}}_{1} \left(\cancel{X} + \frac{Z_{1}}{Z_{123}} \cancel{X} + \frac{Z_{1}+Z_{3}}{Z_{113}} \cancel{X} \right) + m^{2} \cancel{\mathcal{V}}_{1} \right]$$

$$-2m \left(2 l_{11} + 2 \frac{Z_{11}}{Z_{113}} p_{11} + \frac{Z_{1}-Z_{2}+Z_{3}}{Z_{123}} q_{11} \right) \right] \mathcal{U}(P)$$

The terms odd in I will integrate to φ , so we can drop them $= -2e^2 \int_{0}^{\infty} \frac{1}{2} \frac{1}$

-2m[2 =] + Z1-Z2+23 gu]] UCP)
We need to simplify the numerator structure keeping in mind the Gordon Eduntity
between on-shell spinors

$$\widetilde{\mathcal{U}}(P')$$
 $\widetilde{\mathcal{Y}}_{n}$ $\mathcal{U}(P) = \widetilde{\mathcal{U}}(P') \left[\frac{P'_{n} + P_{n}}{2m} + i \underbrace{\mathcal{T}_{n} \cdot Q^{*}}_{2m} \right] \mathcal{U}(P)$

$$\begin{split}
&= \chi l^{\alpha} \, \delta_{\mu} \, \delta_{\nu} \\
&= \chi l^{\alpha} \, \left(2 \, \delta_{\mu \alpha} - V_{\alpha} \, \delta_{\mu} \right) \\
&= \chi \left(2 \, l_{\mu} - \chi \, \delta_{\mu} \right) \\
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&= \chi \left(2 \, l_{\mu} - \chi \, \delta_{\mu} \right) \\
&= \chi \left(2 \, l_{\mu} - \chi \, \delta_{\mu} \right) \\
&= \chi$$

$$\left(\frac{Z_{1}}{Z_{123}} - \frac{Z_{2}}{Z_{123}} + \frac{Z_{1}}{Z_{123}} \right) Y_{1} \left(\frac{Z_{1}}{Z_{123}} + \frac{Z_{1}+Z_{3}}{Z_{123}} \right)$$

We have $\vec{A} \cdot \vec{X} = (\vec{P} - \vec{P}) \cdot \vec{Y}_{n}$ $= M \cdot \vec{Y}_{n} - \vec{P} \cdot \vec{Y}_{n}$ $= M \cdot \vec{Y}_{n} - 2 \cdot \vec{P}_{n} + \vec{Y}_{n} \cdot \vec{P}$ $= 2 \cdot m \cdot \vec{Y}_{n} - 2 \cdot \vec{P}_{n}$ $= 2 \cdot m \cdot \vec{Y}_{n} - 2 \cdot \vec{P}_{n}$ $= 2 \cdot p_{n}' - 2 \cdot m \cdot \vec{Y}_{n}$

also, between spinors,
ASM = -928m

$$= \sqrt[2]{\left[\frac{z_1+z_2}{z_{11}z_3} \frac{z_1+z_3}{z_{12}z_3} q^2 - \frac{3z_1^2+2z_1(z_1+z_3)}{z_{12}z_3} m^2\right] + 2m\frac{z_1}{z_{12}^2z_3} \left[z_{12} + z_{13} + z$$

This gives us
$$\frac{1}{2} \left\{ \frac{1}{2} \right\} \right\} \right\} \right\} \right\} + \frac{1}{2} \left\{ \frac{1}{2} \left(\frac{$$

Notice, the term proportional to $q_n = p_1 - p_1$ integrates to zero as it is odd under the interchange of $Z_2 \longleftrightarrow Z_3$, while the integration measure is even. We can use the Gordon Identify to identify the contribution to F_2 . As we are only interested in this correction, we can ignore the vest

$$= \frac{1}{2m} \frac{1}{2m$$

The gaussian integral is easy (after Wirk to fation) $\int \frac{d^4l}{2\pi^4l} e^{i\frac{\pi}{2}} e^{i\frac{\pi}{2}} = i \int \frac{d^4l}{(2\pi)^4l} e^{-\frac{\pi}{2}} e^{-\frac{\pi}{2}} = \frac{i}{(2\pi)^4l} \left(\frac{\pi}{2}\right)^4$ $= \frac{i}{16\pi^2} \frac{1}{2\pi^2}$

$$\overline{\mathcal{U}(p')} \underbrace{i \mathcal{T}_{uv} q^{v}}_{2m} SF_{z} U(p) = + \underbrace{2ie^{2}m}_{16\pi^{2}} \underbrace{\int_{0}^{2} dz_{1} dz_{2} dz_{3}}_{0} \underbrace{\frac{Z_{1}(Z_{2}+Z_{3})}{(Z_{1}+Z_{2}+Z_{3})^{4}}}_{e} \underbrace{-i \underbrace{\frac{Z_{2}Z_{3}}{Z_{123}}}_{[Z_{123}]} q^{2} + \underbrace{\frac{(Z_{2}+Z_{3})^{2}}{Z_{123}}}_{[Z_{123}]} m^{2}}_{e}$$

× u(ρ) ionr9/mu(ρ)

We are further only interested in the value at q²=0.

Notice the oscillatory nature of the exponential. We could make the change m² > m²-iε, recalling how the is appears in the propagator. This leaves us with

$$\frac{\mathcal{U}(p')}{2m} \frac{1}{2m} \int_{0}^{\infty} \int_{0}^{\infty} \frac{1}{2m} \int_{0}^{\infty} \frac{1}{2m} \int_{0}^{\infty} \frac{1}{2m} \frac{1}{2m} \int_{0}^{\infty} \frac{1}{2m} \frac{1}{2m$$

$$\Rightarrow SF_{2}(0) = + \frac{1}{12} \frac{m^{2} d}{d^{2} d^{2} d^{2$$

rescale all Zi - 1 Zi

$$= +i m^{2} \frac{\chi}{\pi} \int_{0}^{\infty} dz_{1} dz_{2} dz_{3} = \frac{Z_{1}(Z_{23})}{Z_{123}} \int_{0}^{\infty} \frac{Z_{1}(Z_{2$$

$$\delta \vec{F}_{2} = + \frac{\alpha}{\pi} \int_{0}^{\infty} dz_{3} \int_{0}^{\infty} \int_{0}^{\infty} dz_{1} \delta(1-z_{123}) \frac{z_{1}}{z_{23} z_{123}^{4}}$$

$$= + \frac{\alpha}{\pi} \int_{0}^{1} dz_{3} \int_{0}^{1-z_{3}} dz_{2} \frac{1-z_{2}-z_{3}}{z_{2}+z_{3}}$$

Let us also use the Feynman Parameter trick - it's always good to have multiple ways to compute things

$$\overline{u}(p) S \Gamma_{n} u(p) = 2ie^{2} \int d^{4}t \ \overline{u}(p) \left[k y_{n} k' + m^{2} y_{n} - 2m (k_{n} + k'_{n}) \right] u(p) \\
\times \int_{0}^{1} \int_{0}^{1-x} \frac{2}{\left[x(k-p)^{2} + y(k'^{2} - m^{2}) + (1-x-y)(k^{2} - m^{2}) + i\epsilon \right]^{3}}$$

where we have already made use of the S(1-x-y-7) to perform the d7 integral. We want to simplify the denominator by completing the square

$$= (k - xp + yq)^{2} - x^{2}p^{2} - y^{2}q^{2} + 2xy p^{2}q^{2} + yq^{2} - m^{2}(1-2x) + i\epsilon$$

$$= l^{2} - y(y-1)q^{2} + 2xy p^{2}q - m^{2}(1-x)^{2} + i\epsilon$$

$$2p_{1}q_{1} = (p+q)^{2} - p^{2} - q^{2}$$

$$= p^{2} - p^{2} - q^{2}$$

$$= -q^{2}$$

, 92 (0 for the problem we are solving

We need to shift the numerator also

$$k = l + xp - yq \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \frac{\int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{4}} \int_{\mathbb{R}^{3}}^{d^{$$

D = y(1-x-y)(-q2) +m2(1-x0)2-18

= 41e2 fdxfdy fd41 Tup)[xx,x+(xx-yq)x,(xx+(1-yx)+m2x,-2m(2xp+q,(1-2y))]u(p) 12-D]3

Now we need to manipulate the numeroster structure to isolate the term of interest.

$$XY_{n}X = Ll^{2}Y_{n}Y_{d}$$

$$= Ll^{2}(29_{n}u - Y_{d}Y_{n})$$

$$= 2Ll_{n} - LLY_{n} :$$

$$= -\frac{1}{2}l^{2}Y_{n}$$

We want terms proportional to Jurg

but we have to be carreful, as

The Gordon Identity relates

N(p) Y N(p) = N(p) Pu+Pu + i Ju 9 (1) So we also have to find the terms proportional to P,P' We can also use \$ u(p) = m u(p) · (p) が = mu(p)

$$\begin{array}{c} + \sum_{i=1}^{n} \left[\sum_{i=1}^{n} \sum_{i=$$

$$A p + B p' = A'(p'+p) + B'(p'-p)$$

 $= p'(A'+B') + p'(A'-B')$
 $A = A' \cdot B'$
 $B = A' \cdot B'$
 $A' = \frac{1}{2}(A+B)$
 $B' = \frac{1}{2}(B-A)$

$$\widetilde{u}(p) \, \delta \, \widetilde{u}(p) = 4 \, i \, e^2 \, \int_0^1 dx \, \int_0^{1-x} \int_0^1 \frac{d^4 p}{(2\pi)^4 1} \, \frac{1}{[2^2 - 2 + i \, e]^2}$$

$$\times \, \widetilde{u}(p) \, \left\{ \, \mathcal{Y}_{n} \left[-\frac{1}{2} \, \ell^2 + m^2 \left(1 - X \right)^2 + q^2 \left(x + y \right) (1 - y) \right] \right.$$

$$+ 2 \, m \left(p'_{n} + p_{n} \right) \, - \frac{1}{2} \, x \, (x - 1)$$

$$+ 2 \, m \left(p'_{n} - p_{n} \right) \, \frac{1}{2} \, \left(2 - x \right) \left(2 y - 1 + x \right) \, \left[1 \, u(p) \right]$$

We can use the Gordon Identity to preplace

As we are interested in the correction proportional to iturgi,

we can just focus on this terms.

$$\Rightarrow \delta \overline{I_{2}} = 4ie^{2} \int_{0}^{1} dx \int_{0}^{1-x} dy \int_{0}^{1-x} \frac{1}{[2^{2}-\Delta+i\epsilon]^{3}} (-zm^{2}) \times (x-1)$$

$$= +8im^{2}e^{2} \int_{0}^{1} dx \times (1-x) \int_{0}^{1-x} \int_{0}^{1-x} \frac{1}{[2^{2}-\Delta+i\epsilon]^{3}}$$

$$\frac{-i}{(4\pi)^{2}} \frac{1}{2} \frac{1}{\Delta}, \quad \Delta^{*} = m^{2}(1-x)^{2} - q^{2} \times (1-x-y)$$

=
$$m_3 \frac{\pi}{\alpha} \int_0^0 \sqrt{(1-x)} \int_{1-x}^0 \frac{m_3(1-x)_3 - d_3 \lambda(1-x-3)}{1}$$

This integral is finite and four of IR(q2=0) divergences. It is related to a log (arcCotu). We will just focus on the q2=0 part

$$=$$
 $y-2=28F_2=\frac{4}{10}$

Why is g-2 fruit UV divergences @ + loop?

- Simply stated, the operator that would be needed to absorb a UV divergence would be

to have the cornect Lorentz structure to be able to carried divergence. However, assuming QED is a renormalizable QFT, we know this operator is not allowed as it is dimension - 5. Without such an operator, there can be no UV divergence, or the theory is not correct or not "renormalizable".

Why is g-2 from of IR divergence @ 1-loop?

henerically, IR divergences are associated with long-range physics. Because we can undustand the classical world with classical physics, with our QFT description of the world, there can not be such an IR divergence, else our classical EM description would not work.

Specifically for g-2 @ Hoop in QED, we can show with perturbation theory, there can not be an IR divergence for g-2. Ih divergences asise from "soft photon radiation."

At LO in QED, the IR divergence comes from two graphs (in addition to vertex correction)

lets look at the Lorentz structure of the first term

The IR divergence will come when k becomes

co-linear with p' or p (p.k=0 or p'k=0)

iA = U(p') Yu i Yv Exect) U(p)

 $k_mu = Q(1,0,0,1)$ so we can also let $Q \rightarrow 0$ to get the divergence

= 1 u(pi) 8u(x-1+m) 2* u(p) (p-k)2-m2 tie

8 x = - 2 x +2 E.P 7 =0

= 1 m(p) / [-47 +2E-p - 128* +m9*] u(p)

= i u(p') [1/2 E* p + 1/2 E* k] u(p) As k > 0", only the

first term is IR divergent.

However, this term is only proportional to rep Yullp) and so does not have the comeet Lonentz structure to contribute to g-2. The same holds for the other graphs